

Modelling the climate of the last millennium in the LSCE : the ESCARSEL project



Eruption solaire

CLIM (modélisation du CLIMat) & **ESTIMR**
(Extrêmes, STatistiques et Modélisation Régionale)
teams



Pinatubo 1991

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Questions :

Climate variability on the last millennium:

-> **Medieval Warm Period**

-> **Little Ice Age**

-> How (where and which amplitude), and causes

Climate models used to study the last millennium for the AR4:

- GCM = 6/12

- Intermediate complexity = 5/12

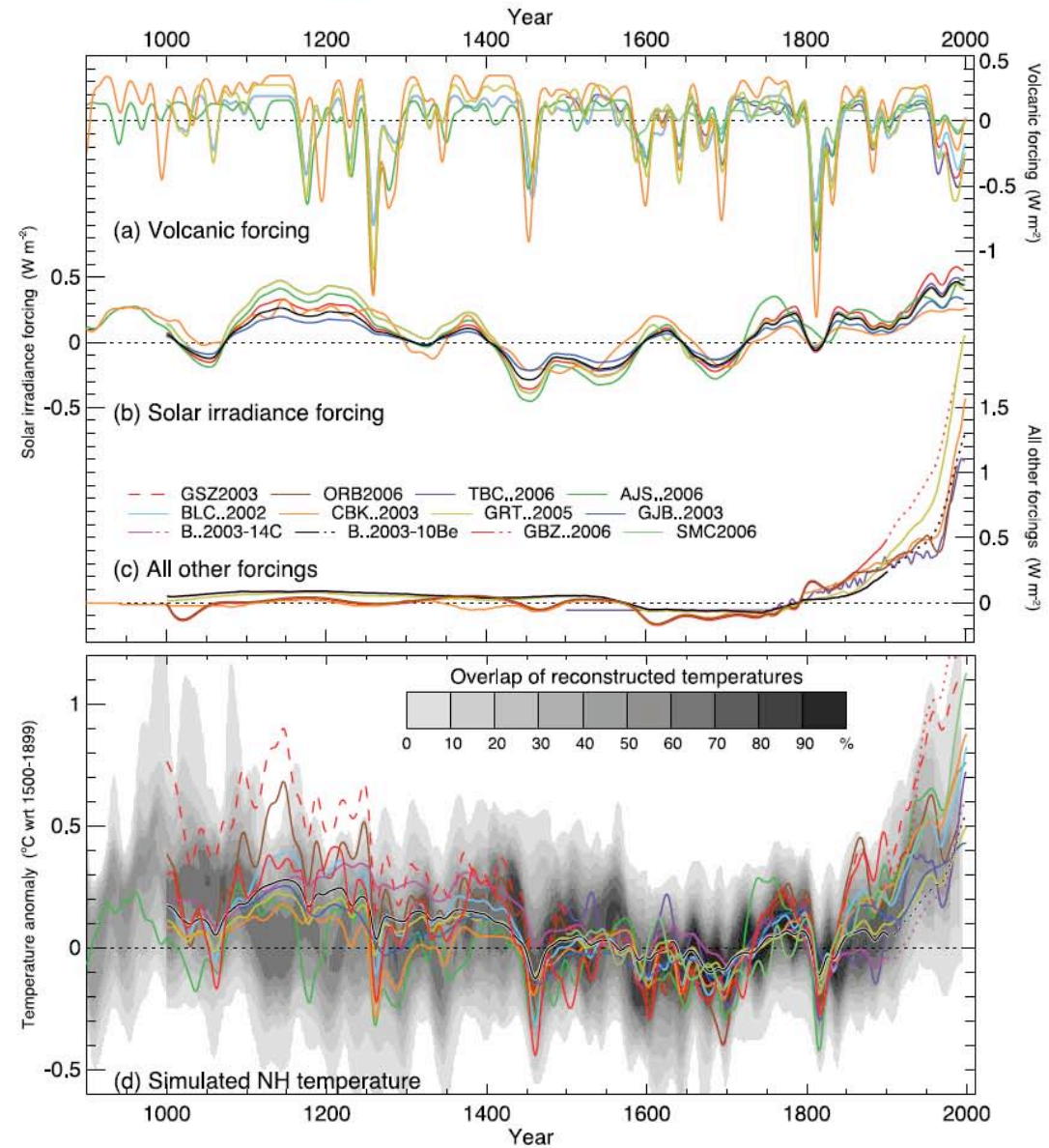
- EBM = 1/12

Forcings :

- Solar irradiance

- volcanic aerosols

- land use (too many uncertainties)



IPCC AR4, chapt 6

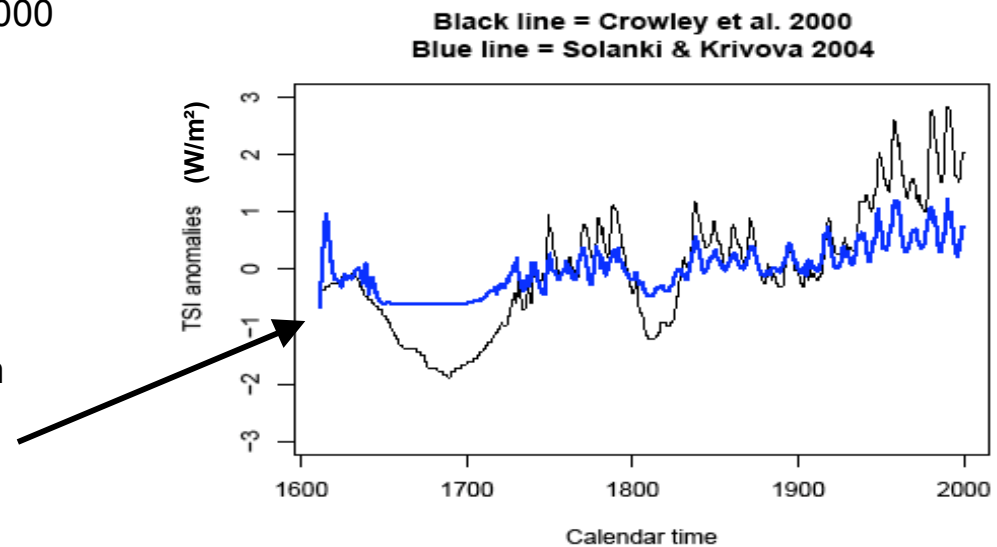
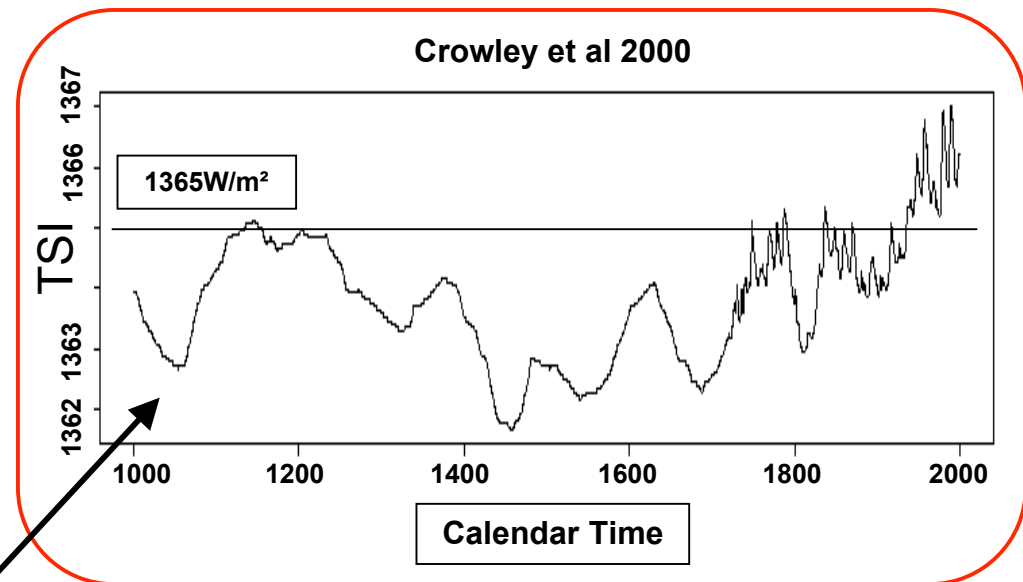
Two main forcings : solar and volcanic

Models and simulations

- ISPLCM4v2 96 / 72 / 19
- 1 CTRL, 1 solar run and 1 volcanic run
- CNRM CM3 T63 / L31
- 1 CTRL and 1 simulation with both solar and volcanic forcing

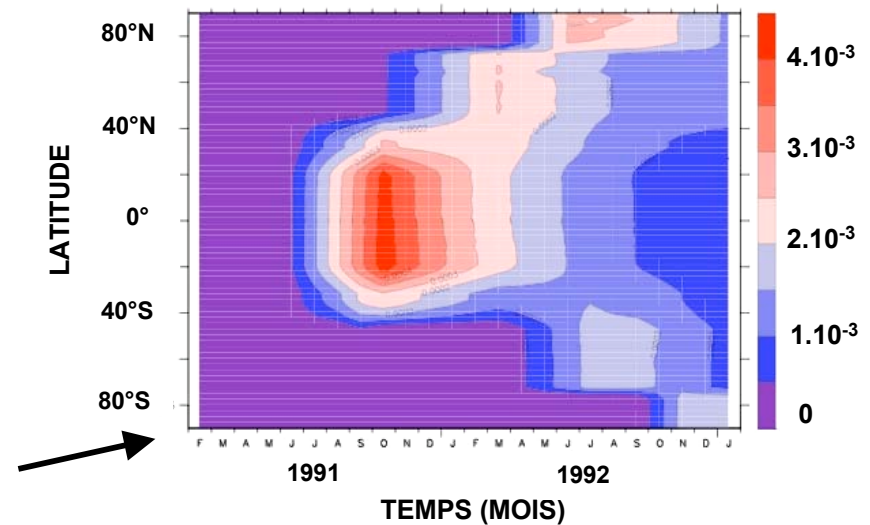
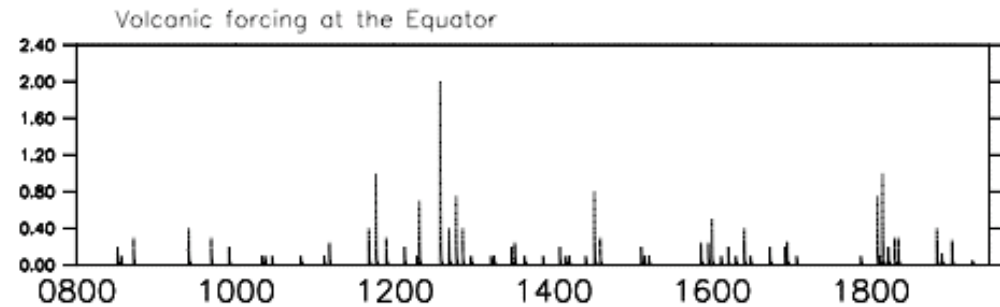
Solar forcing

- Crowley et al 2000 (based on Bard et al 2000 & Lean et al 1995)
- anomalies calculated from the 1850-2000 mean (ENSEMBLES)
- - **0.25%** of the TSI for the Spörer minimum
- Solar variability = 0.1% TSI ?



Volcanic forcing

- IPCC AR4 = tuning of the solar irradiance
- 2D Reconstitution (latitude-time) of the aerosols optical parameters (Ammann et al 2008) of the main eruptions
- Tropical eruptions ; Volcanic Explosive Index > 4
- Implementation in the radiative code of LMDz (2 levels above the tropopause)



$$I/I_0 = e^{-(\tau)}$$

**Optical thickness (τ) of the volcanic aerosols
From the Mt Pinatubo eruption (DJF 1991-1992)**

Northern Hemisphere Temperature

Global agreement

- Good!
- Missing volcanoes
- Problem 1000-1200

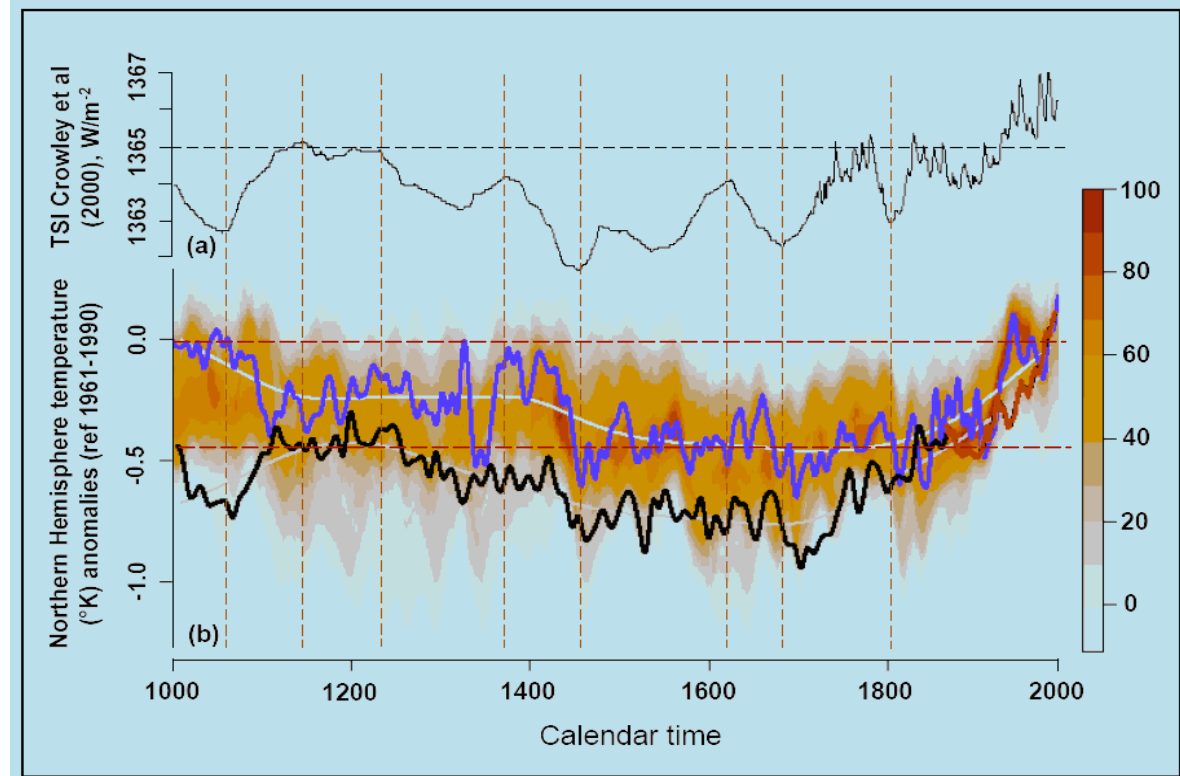
Correlations

- with NH temp reconstructions = between **0.31 & 0.61**

Variability

- Low frequency Max-Min on 1200-1850 = **0.34**
- Standard deviation of the residuals = **0.071**

=> TSI variability = 0.1 or 0.25% ?

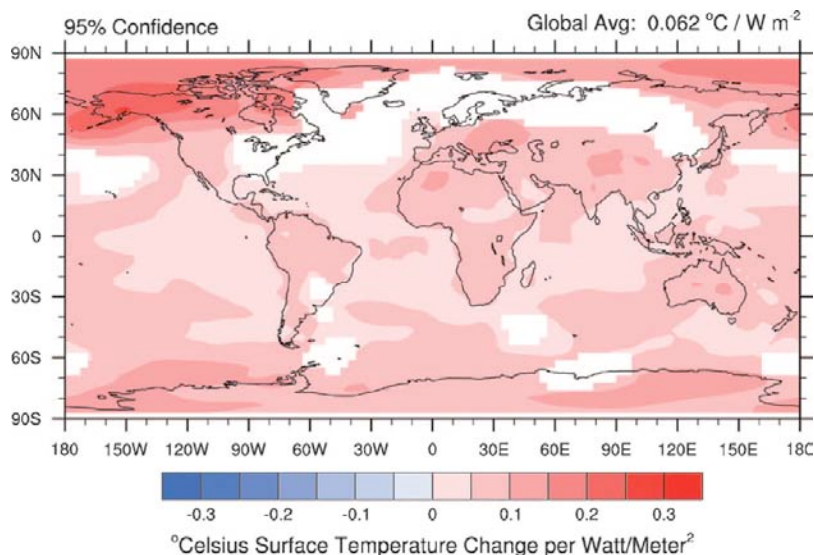


Correlation with solar simulation	Mann et al 2008 (EIV land + ocean)	Moberg et al 2005	Crowley&Lowery 2000	Ammann&Wahl 2007	Preind. CTRL
1001-1850	0.40	0.58	0.52	0.36	-0.1
1201-1850	0.54	0.61	0.50	0.32	-0.05
Residuals (1001-1850)	0.21	0.30	0.1	0.04	-0.12
Residuals (1201-1850)	0.37	0.26	0.15	0.13	-0.07

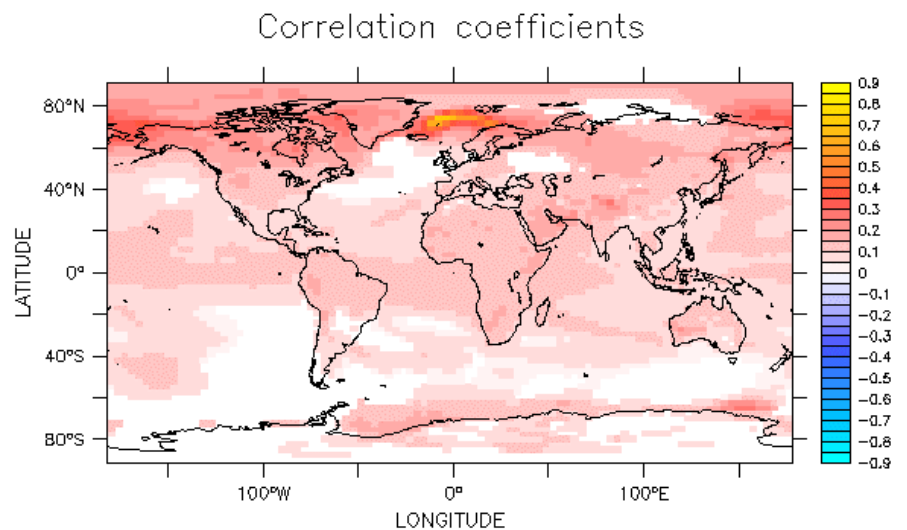
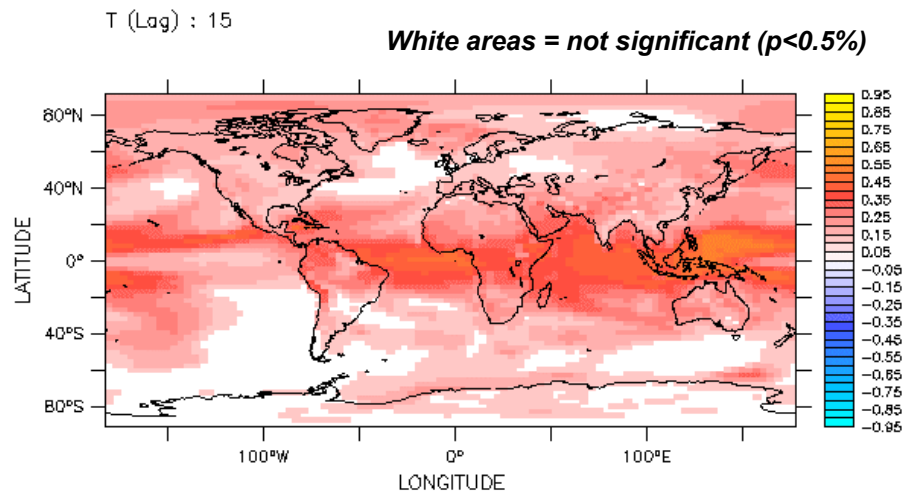
	Solar simulation IPSLCM4v2	Preind. CTRL	Mann et al 2008 (EIV land + ocean)	Moberg et al 2005	Crowley&Lowery 2000	Ammann&Wahl 2007
Standard deviation of the residuals	0.071	0.054	0.106	0.107	0.045	0.067
Range (max-min) of the LF trend (1001-1850)	0.34	0.063	0.46	0.644	0.225	0.156
Range (max-min) of the LF trend (1201-1850)	0.34	0.022	0.23	0.432	0.195	0.135

Regression and correlation between TSI & temperature

- 1000-1800
- annual temperature
- Max with 15yr lag
- Sensitivity coefficient = $0.09^{\circ}\text{K}/\text{Wm}^{-2}$ (lag 0)
- important reaction of the arctic sea ice



Ammann et al. 2007 PNAS
NCAR-CSM1.4 ; T31

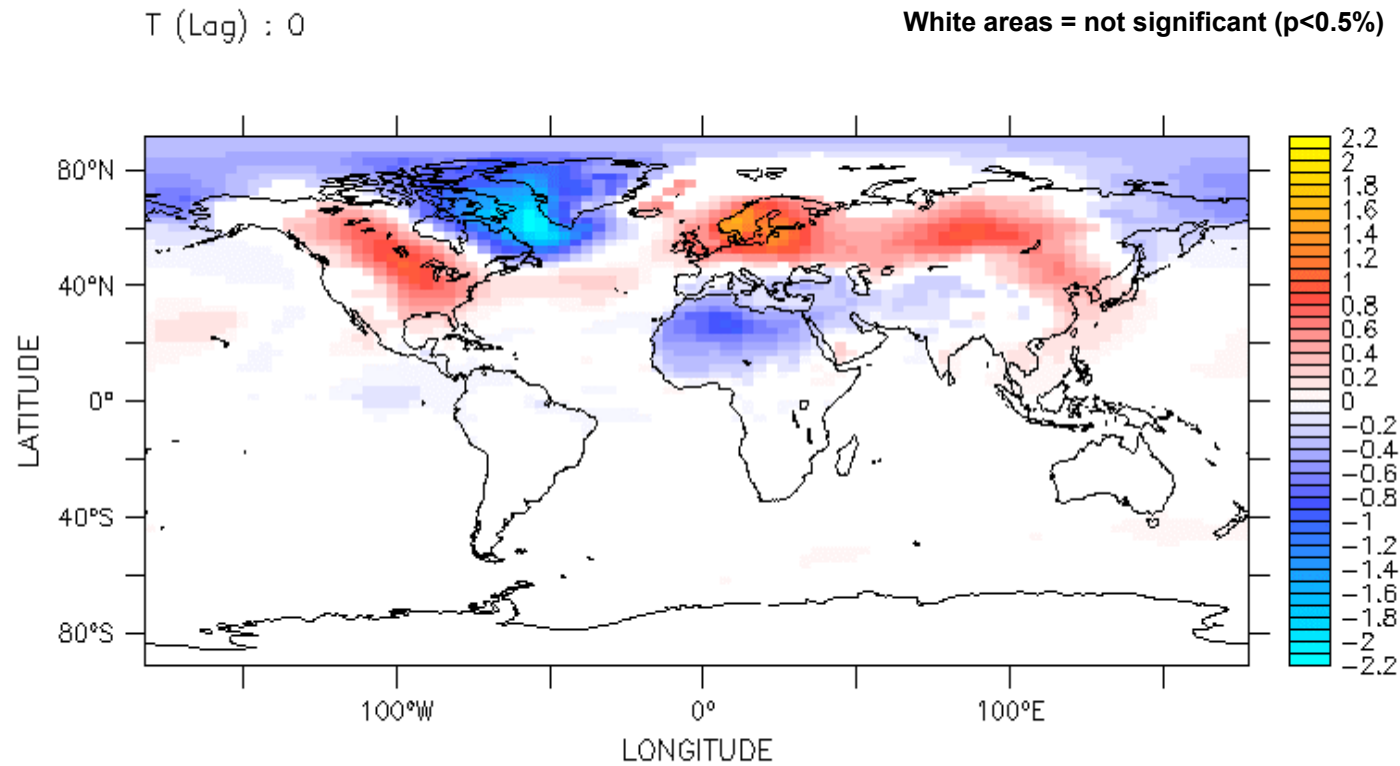


Regression coefficients ($d(t_{2m})/d(TSI_{anom})$)

Solar simulation IPSLCM4v2

Regression t_{2m} vs NAO

- 1000-1800
- NAO pattern (regression $t_{2m}(\text{DJF}) / \text{NAO index}$)



Perspectives

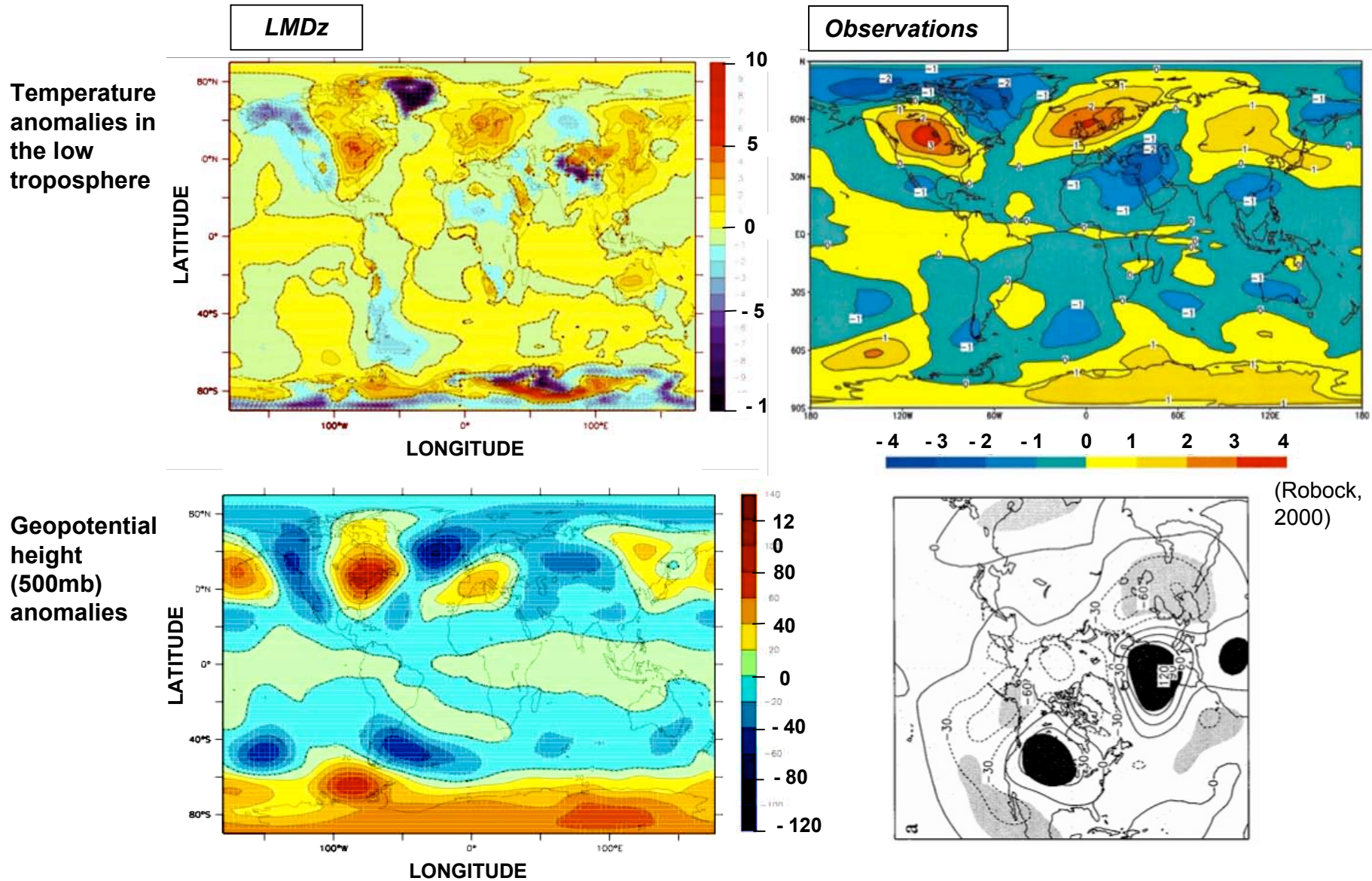
In progress...

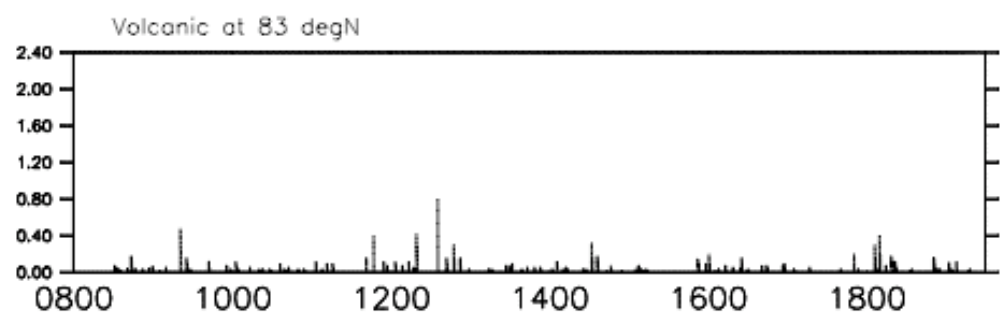
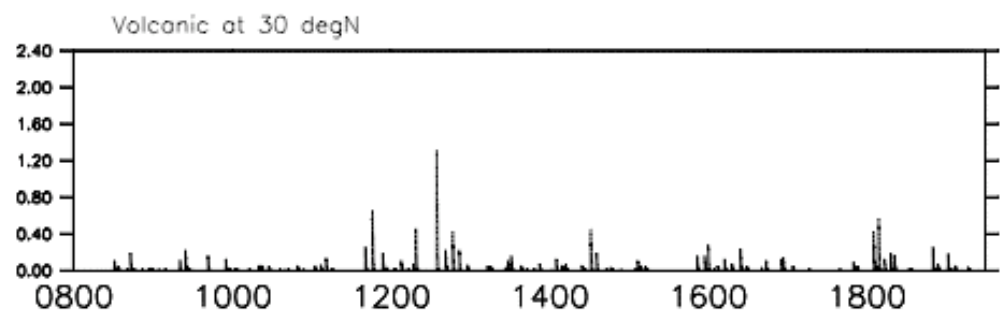
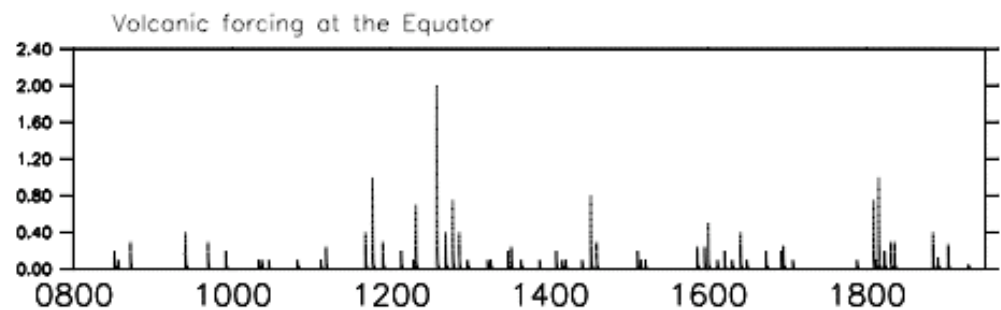
- Model sensitivity to solar forcing
- volcanic simulation

Up to come...

- run with both forcings = solar + volcanic
- ensembles on selected periods (LIA, MWP)

Tests with LMDz on Pinatubo (1991) : results on DJF 1991–1992 (winter after eruption)





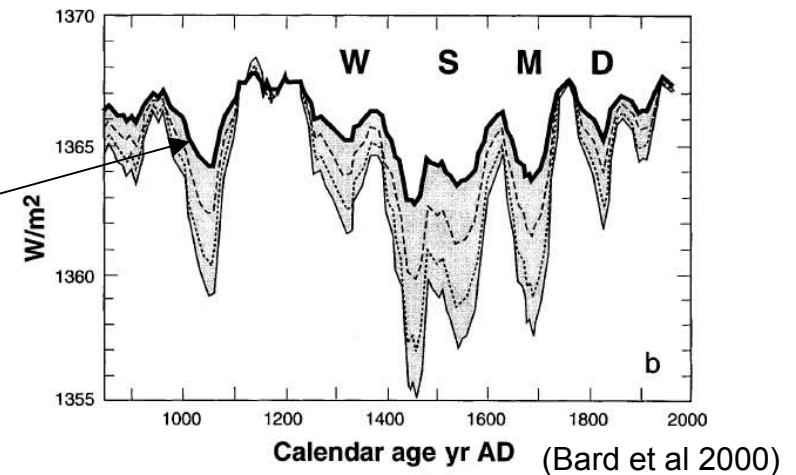
Solar forcing

Variability of the total solar irradiance (TSI)

- **Sun spots ...** : Sum [Sun spots + faculae] > 0 ; 11yr cycle

- Link between the TSI and the solar magnetic activity (Willson & Hudson 1988 ; Willson 1997)
- Reconstitution based on cosmo nucleides (^{10}Be & ^{14}C)

=> **Bard et al. 2000** (scaling = - 0.25% TSI for the Maunder Minimum (Lean et al 1995)



- ... + a multidecadal variability

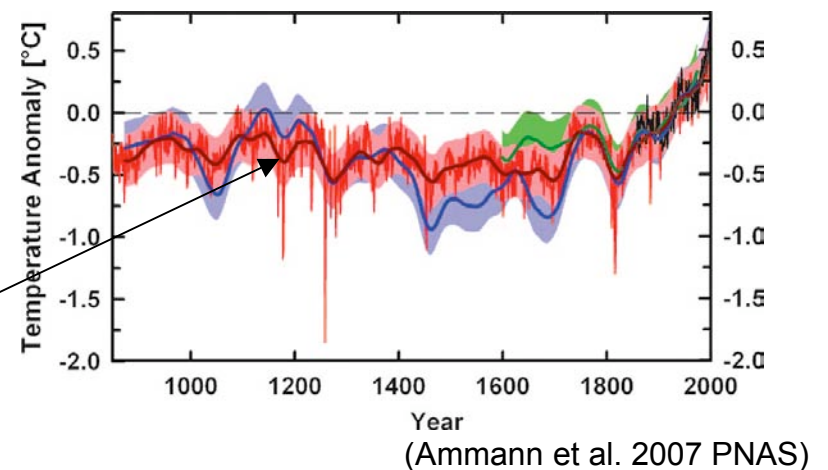
Ammann et al. 2007 = test of different scaling on the TSI reconstruction provided by Bard et al 2000, with a GCM (NCAR)

⇒ - 0.1%

⇒ - 0.25%

⇒ - 0.65%

⇒ **Global temperature** : - 0.25% TSI at the Maunder Minimum

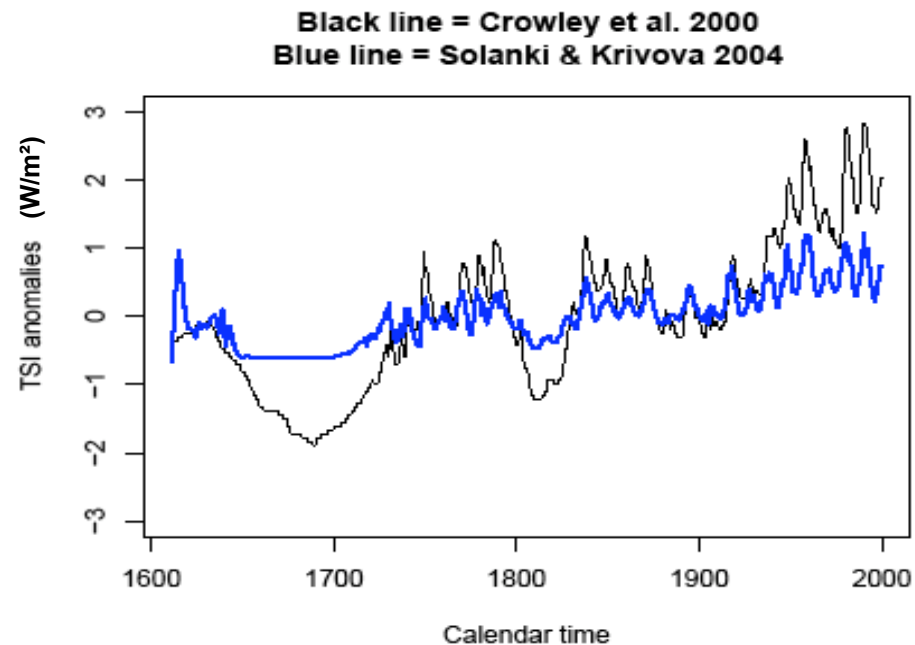
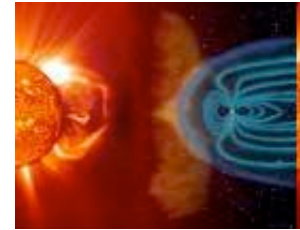


Uncertainties on solar forcing

Is there any secular variability, other than sun spots?

⇒ Last PMIP3 meeting :

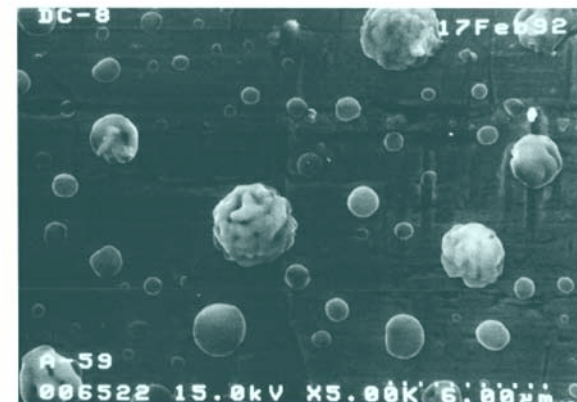
- consensus : secular variability less than 0.1% = only sun spots
- reconstitution provided by Solanki et Krivova 2004
- still an open debate!



Volcanic forcing

- Sulphate aerosols in the stratosphere
- volcanic eruptions with a VEI > 4
- essentially tropical eruptions

Plane view of stratospheric H₂SO₄/H₂O droplets and sulfuric-acid-coated volcanic ash particles collected on February 17, 1992, near (53°N; 67°W) at 10.7 km

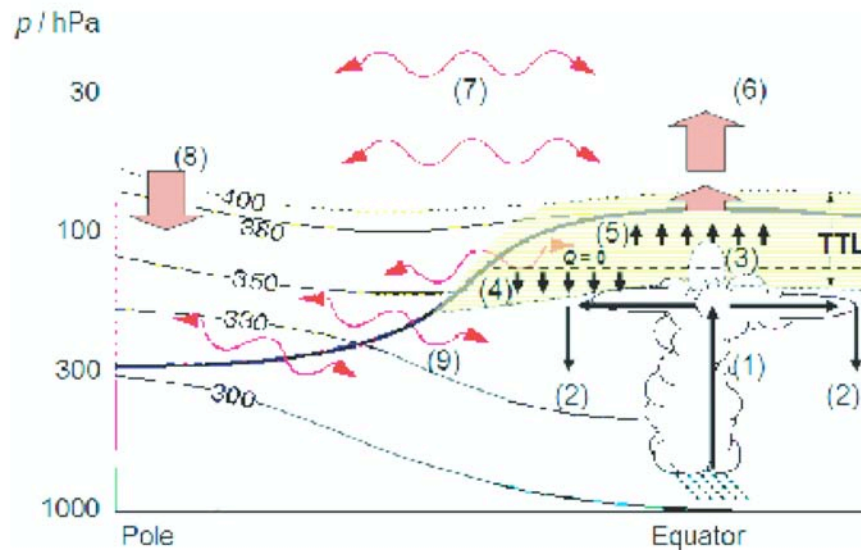


VEI	HAUTEUR DU PANACHE	VOLUME D'ÉJECTION	CLASSIFICATION	EXEMPLE
0	<100 m	1000s m ³	Hawaiano	Kilauea
1	100-1000 m	10000s m ³	Hawaiano/Estromboliano	Stromboli
2	1-5 km	1000000s m ³	Estromboliano/Vulcaniano	Galeras (1992)
3	3-15 km	10000000 m ³	Vulcaniano	Ruiz (1985)
4	10-25 km	100000000s m ³	Vulcaniano/Plineano	Galunggung (1982)
5	>25 km	1 km ³	Plineano	St. Helens (1980)
6	>25 km	10s km ³	Plineano/Ultra-Plineano	Krakatau (1883)
7	>25 km	100s km ³	Ultra-Plineano	Tambora (1815)
8	>25 km	1000s km ³	Ultra-Plineano	Toba (74 ka)

Volcanic eruptions with stratospheric impact : VEI > 4

Transport of stratospheric aerosol

Stratospheric transport



Schematic of the aerosol life cycle (Hamill et al., 1997)

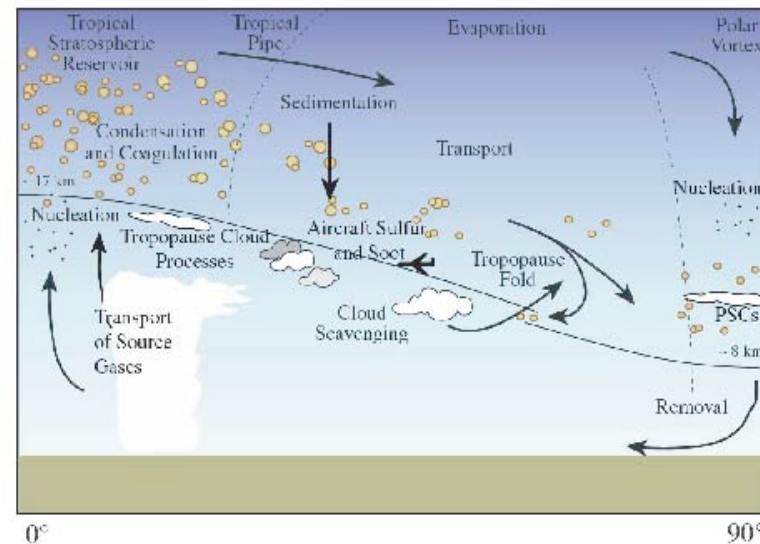


Illustration of major transport processes in the troposphere/stratosphere system. The Tropical Tropopause Layer (TTL) is given in yellow. Black solid line: (cold-point) tropopause. Thin dashed line: zero net radiative heating level ($Q = 0$). Numbers at contours: potential temperature levels. The TTL is between approximately 340–400 K potential temperature. Bracketed numbers: (1) Deep convection lifting low level air into the upper troposphere, horizontal arrows indicating main detrainment. (2) Downward branches of tropospheric Hadley circulation. (3) High cloud tops penetrating TTL. (4) Descent in TTL below $Q = 0$. (5) Radiatively driven ascent in TTL above $Q = 0$. (6) Upward branch of Brewer-Dobson circulation (tropical pipe). (7) Stratospheric stirring and meridional transport by breaking waves. (8) Large scale subsidence. (9) Quasi-adiabatic exchange across tropopause. (Th. Peter, personal communication, modified after Holton et al. [1995].)

Major volcanic eruptions of the past 250 years with stratospheric impact (Pitari et al., 2002)

Volcano	Latitude	Year	VEI	DVI	SO₂ (Tg)	IVI	ΔT (K)
Laki	64 N	1783	4	2300	43–65	0.19	−0.14*
Tambora	8 S	1815	7	3000	34–175	0.50	0.18*
Cosiguina	13 N	1835	5	4000		0.11	−0.32*
Askja	65 N	1875	5	1000		0.01	
Krakatau	6 S	1883	6	1000	15	0.12	−0.25*
Tarawera	38 S	1886	5	800		0.04	
SantaMaria	15 N	1902	6	600		0.05	−0.14*
Ksudach	52 N	1907	5	500		0.02	
Katmai	58 N	1912	6	500	12	0.15	
Agung	8 S	1963	4	800	12	0.06	−0.31*
St.Helens	46 N	1980	5	500	2.1	0.00	
El Chichon	17 N	1982	5	800	3.3–10	0.06	−0.12+
Pinatubo	15 N	1991	6	1000	20	0.13	−0.20+

VEI = volcanic explosivity index; DVI = dust veil index;
IVI = ice core volcanic index; ‘*’ = Northern Hemisphere; ‘+’ = Global

Implémentation de l'impact radiatif des aérosols volcanique dans MDZ

➤ Code de Mie:

Calcul de l'albédo de simple diffusion (ω) et le facteur d'asymétrie (p) pour les aérosols stratosphériques sulfatés en phase aqueuse (forme binaire H₂SO₄/H₂O: 75%/25%)

El Chichon + Pinatubo (SO₄ droplet):

R : rayon modal en nm : **500.0**

Sigma : largeur de distribution: **1.30**

ω_1 ω_2 : albedo de simple diffusion sur les 2 bandes du visible : **1.000** **0.995**

g_1 g_2 : paramètre d'asymétrie sur les 2 bandes du visible : **0.7079** **0.7548**

➤ Implémentation de l'épaisseur optique (Tau) sur les 2 couches atmosphériques au dessus de la tropopause

Epaisseur optique des aérosols volcanique
Pour l'éruption du Mt Pinatubo (1991-1992)

